Space Weather Lecture 1: Introduction



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 Weather is the state of the atmosphere, to the degree that it is hot or cold, wet or dry, calm or stormy, clear or cloudy.

Human activities and technologies have always been prey to the

extremes of weather









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... around the middle of the 19th century

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I. Introduction: Space Weather definition

 Space Weather are disturbances of the upper atmosphere and near-Earth space environment driven by the magnetic activity of the Sun.





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 Space Weather are disturbances of the upper atmosphere and near-Earth space environment driven by the magnetic activity of the Sun.



A planet habitability depends on space weather!

I. Introduction: some of the most recent events

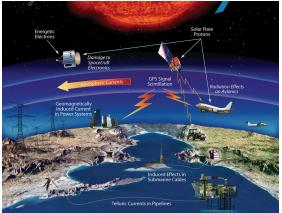
- Capella's Earth-imaging satellites deorbiting faster than anticipated (August 4, 2023)
- 40 from 49 Starlink satellites were burned in the atmosphere (February 4, 2022)
- Northern lights above Zugspitze (September, 25, 2023)
- Mother's Day super storm (May, 10, 2024)

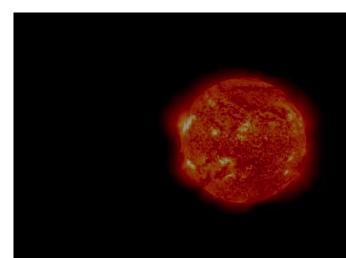




Credit: Ralf Plechinger

 Since middle of the 19th century growth of the electric power industry, the development of telephone, radio, space-based communications and navigation systems has dramatically increased the vulnerability of modern society to space weather.





I. Motivation: to learn about

- the potential societal and economic impacts
- the technologies used for forecasting Space Weather events
- the infrastructure behind Space Weather services
- the physics behind Space Weather phenomena
- \bullet \sim 9 lectures and 2 tutorials



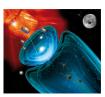


Credit: NOAA SWPC

Gredit: W. Redal

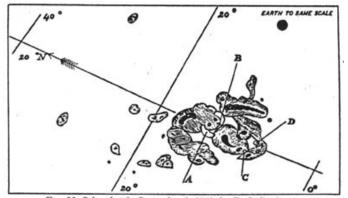
Lecture's programm

- Introduction (What is space weather? The potential societal and economic impacts, the technologies used for forecasting Space Weather events, the infrastructure behind Space Weather services)
- The Sun and the Solar Wind
- Bow Shock
- Kelvin-Helmholtz Instability and Field Line Resonances
- Earth's Magnetic Field
- Interaction with Interplanetary Magnetic Field and Reconnection
- Magnetospheric Substorms and Storms
- Radiation belts
- Ionosphere



II. Historical Background: The Carrington Event August-September 1859

 Richard Carrington (England) noted an outburst of "two patches of intensely bright and white light" from a large group of sunspots near the center of the Sun's disk – solar flare



Richard Carrington's 1859 F10. 36. Solar sketch, September 1, 1859, by R. C. Carrington drawing

II. Historical Background: The Carrington Event August-September 1859

 It was followed the next day by auroras seen e.g. in Sub-Saharan Africa, Mexico, Cuba, Hawaii, and even in Colombia.





Church's 1865 painting "Aurora Borealis"

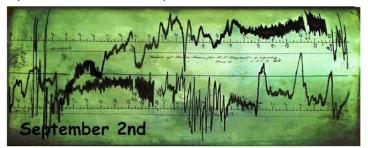
"The Aurora Borealis, seen from the pier, Boulogne, 1853"

II. Historical Background: The Carrington Event August-September 1859

 Shortly after midnight on September 2, 1859, campers in the Rocky Mountains were awakened by an "auroral light, so bright that one could easily read common print. Some of them partly insisted that it was daylight and began preparation of breakfast." (The Rocky Mountain News)

II. Historical Background: The Carrington Event August-September 1859

- A magnetic storm was also observed. The storm strength range (Dst) from -800 nT to -1750 nT. (March 14, 1989, -589 nT; November 20, 2003, -422 nT; May 10, 2024, -412 nT)
- "The solar spots, the mean daily range of the magnetic needle and the frequency of auroras – are somehow dependent the one upon the other". (Elias Loomis, US, 1860)



A magnetogram Greenwich Observatory, Declination, or compass direction, (D) is the lower trace and the horizontal force (H) is the upper trace.

II. Historical Background: The Carrington Event Socioeconomic impacts

- Disruptions of telegraph service "at the busy season when the telegraph is more than usually required" (Walker, 1861),
- the telegraph companies associated loss of income.



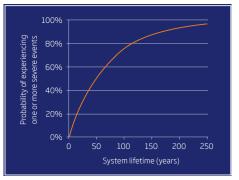
A STORM OF ELECTRICITY

TELEGRAPH WIRES USELESS FOR SEVERAL HOURS.

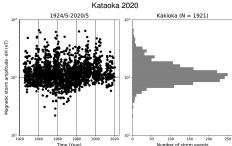
ONE OF THE MOST SEVERE DISTURBANCES FOR MANY YEARS, EXTENDING EVEN TO EUROPE—TELEPHONE WIRES ALSO OB-STRUCTED—BUSINESS DELAYED A GOOD PART OF THE DAY.

Yesterday's storm was accompanied by a more serious electrical disturbance than has been known for years. It very seriously affected the workings of the telegraph lines both on the land and in the sea, and for three hours—from 9 A. M. until noon—telegraph business east of the Missisippi and north of Washington was at a stand-still.

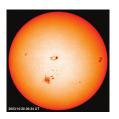
II. Historical Background: Probability of the Carrington Event

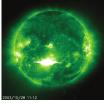


Estimated economic impact of loss of power ~M€9,344.04 direct, taken from "Extreme space weather: impacts on engineered systems and infrastructure", Royal Academy of Engineering



II. Recent event: Magnetic Storm October, 2003



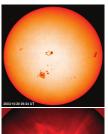


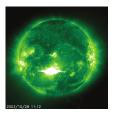
 A large sunspot (SOHO/MDI image, upper-left) erupted with a strong x-ray flare (SOHO/EIT image, upper-right).

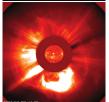
SOHO spacecraft observations in a halo orbit around the Sun–Earth L1 point.

Credit: NASA

II. Recent event: Magnetic Storm October, 2003





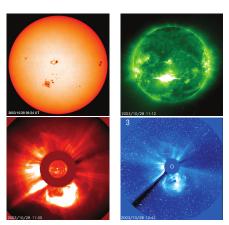


SOHO spacecraft observations in a halo orbit around the Sun–Earth L1 point.

Credit: NASA

- A large sunspot (SOHO/MDI image, upper-left) erupted with a strong x-ray flare (SOHO/EIT image, upper-right).
- Within minutes, SOHO/LASCO detected a halo coronal mass ejection (CME) emerging from the Sun (lower-left).

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- Within minutes, SOHO/LASCO detected a halo coronal mass ejection (CME) emerging from the Sun (lower-left).
- 1.5 hour after the flare, a shower of energetic protons reached the SOHO spacecraft, creating the "snow" in the lower-right image.
- This CME triggered powerful magnetic storm (~-400 nT).

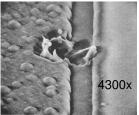
III. Effects of extreme space weather on modern technology

can be devided in several categories:

- (a) Power grid outages
- (b) Interference with Global Navigation Satellite System (GNSS) signals
- (c) High Frequency (HF) radio communication blackouts
- (d) Spacecraft hardware damages







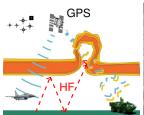
Credit: US Air Force Research Laboratory (AFRL)

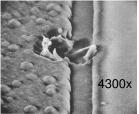
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Credit: US Air Force Research Laboratory (AFRL)

III.a Effects: Power Grids – Examples

 The March 13, 1989 magnetic superstorm: a voltage depression on the Hydro-Quebec power system in Canada. The system collapsed within seconds. The province Quebec was blacked out for approximately 9 hours.

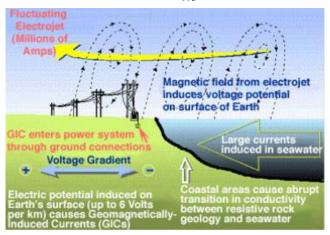




Hydro Quebec, Electric Power Transformer

III.a Effects: Power Grids - Reason

- Power transmission systems are vulnerable to Direct Currents (DC) driven by magnetic induction
- Induced electric potential: $e = -\partial_t \iint_S \mathbf{B} \cdot \mathbf{dS}$



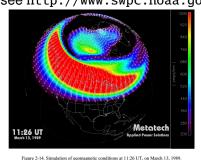
III.a Effects: Power Grids – Reason

 Too much DC current causes hot spots, wires melt and oil baths catch fire



III.a Effects: Power Grids – Solutions?

- Operators of the North American power grid constantly analyze potential risks associated with space weather events
 - * monitoring voltages and ground currents in real time
 - in case of significant currents invoke conservative and mitigating operation practices.
- Forecasts are produced by National Oceanic and Atmospheric Administration's (NOAA) Space Weather Prediction Center (SWPC): see http://www.swpc.noaa.gov





III.b Effects: GNSS

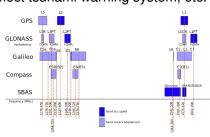
GNSS Services

- The GNSS is a constellation of satellites which provides global coverage.
- Combined data product (4 times hourly) had better than 10 cm accuracy during the October 2003 magnetic storm.
- However, high-rate and real-time GNSS analysis can be critical, e.g. in aviation, autonomic driving, in detecting seismic surface waves, as a consequence affect tsunami warning system, etc.

GLONASS – Global Navigation Satellite System (Russia)

COMPASS – BeiDou Navigation Satellite System (China)

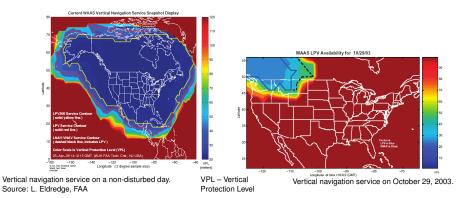
SBAS – satellite-based augmentation system (data from GNSS satellites is analysed on the ground and sent to users)



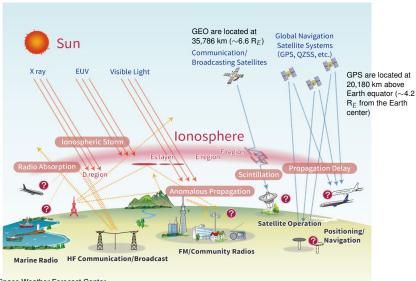
III.b Effects: GPS - Example

Aviation Navigation

GPS was not available for 30 h in Oct 2003 leading to flight delays.



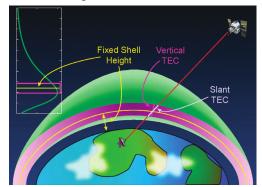
III.b Effects: GNSS - Reason



Source: Space Weather Forecast Center

III.b Effects: GNSS - Reason

- The ionospheric corrections are done using a thin shell model.
- The accuracy depends on the Total Electron Content (TEC) in the ionosphere.
- During significant ionospheric disturbance the model can be used only for the horizontal navigation but not for the vertical.



Source: L. Eldredge, FAA

III.b Effects: GNSS - Solutions?

Aviation Navigation

- To avoid such problems, switch to dual-frequency GNSS system has been started (accuracy increase from ~meters to ~centimeters).
- In May 2018, Xiaomi launched the first dual-frequency GNSS smartphone.
- However, during high ionospheric disturbances the signal is still inaccurate (can be \sim 50m).
- Build better ionospheric models



III.c Radio blackout – Example

May 23, 1967

- Radar system designed to detect incoming Soviet missiles was disrupted, in what the military perceived to be an act of war.
- US Air Force authorized nuclear missile-carrying aircraft.
- Information from space-weather forecasters, who realized that it was a solar flare jamming the radar, managed to prevent military action.



Three radar stations at the Ballistic Missile Early Warning System in Anderson, Alaska, in 1962. Image: Wikimedia



The solar flare begins at exactly 18:40 UT on May 23, 1967. Image: National Solar Observatory

III.c Radio blackout - Example

USS Midway aircraft carrier: crewman's story

- The HF communication was occasionally interrupted by solar flares
- Could not do anything: needed to wait



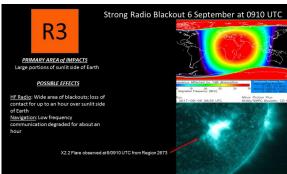


Source: Wikipedia

III.c Radio blackout - Reason

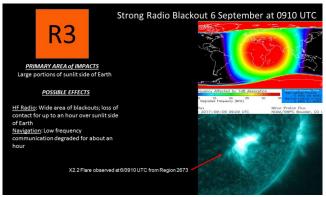
The latest Space Weather event: Sept 6, 2017

- An active region on the sun belched out two huge streams of radiation. One of them was the largest such flare in over a decade.
- These two flares were placed in the X-class, the most powerful type of solar flare. X-class solar flares are the largest explosions in the solar system.



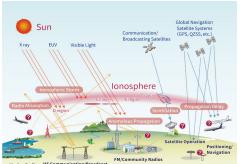
III.c Radio blackout - Reason

- The sudden outburst of electromagnetic energy travels at the speed of light. Therefore any effect is observed at the same time of the event.
- Increased level of X-ray and extreme ultraviolet radiation results in ionization in lower layers of the ionosphere on the sunlit side of Earth.



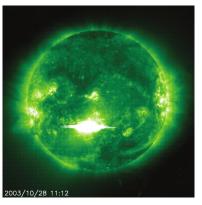
III.c Radio blackout - Reason

- A strong enough solar flare produces ionization in the lower, more dense layers of the ionosphere.
- Radio waves that interact with electrons in layers lose energy due to more frequent collisions in the higher density environment. This can cause HF radio signals to become degraded or completely absorbed.
- This results in a radio blackout the absence of HF communication, primarily impacting the 3 to 30 MHz band.



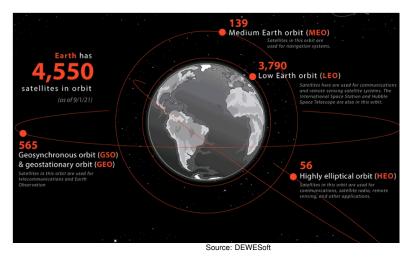
III.c Radio blackout - Solutions?

• Learn how to predict solar flares, e.g., using artificial intelligence



III.d Effects: Satellites

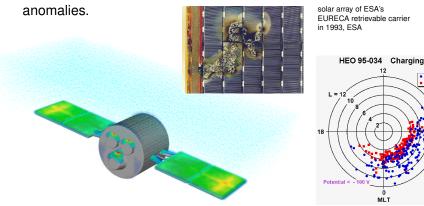
• The current fleet of active satellites orbiting the Earth is \sim 9100 (August 2024, Source: Statista).



III.d Effects: Satellites - Examples

• On January 20, 1994, Telesat Antik E1 and E2 were disabled for \simeq 7 hours. Canadian press, TV and data services were lost.

The electrostatic discharge is one of the major causes of spacecraft

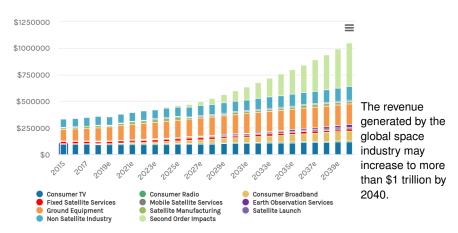


Fennell+08, several years of observations

• L - upper

III.d Effects: Satellites - Investing in Space

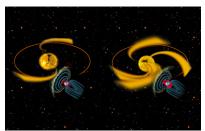
The Global Space Economy (\$t)



Source: Haver Analytics, Morgan Stanley Research forecasts

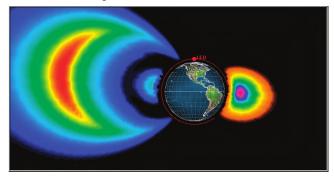
III.d Effects: Satellites - Reasons

- The satellites are damaged by enhancements of the magnetospheric electron intensity. Most of anomalies are related to major magnetic storms (like in 2003).
- They are also associated with high-speed streams emanating from the coronal holes during declining phase of the solar cycle.
- Also when storm and high-speed streams are absent the damages can occur. During substorms, injected energetic plasma into the inner magnetosphere can cause electrical charge to build up on spacecraft surfaces. The electrostatic discharge occurs subsequently.



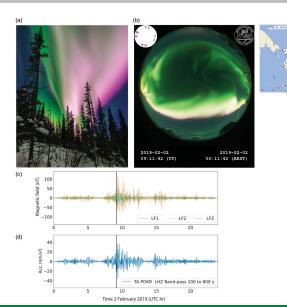
III.d Effects: Satellites - Solutions?

- To create more accurate long-term models of the radiation belts and ring current
- To better observe damages and then build better satellites



Credit: D. Chenette

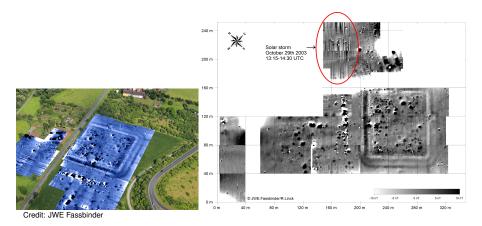
III. Effects: Related to Seismology



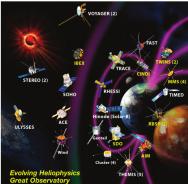
Credit: Tape+2020

III. Effects: Related to Archeology

- Archaeological site (the Roman fort Wörth am Main) overlaid by a magnetogram.
- The magnetic storm has produced stripes in the magnetograms.

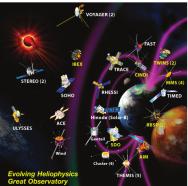


 Although heliospheric missions are all primarily for scientific research, they provide much of space weather data used by both civilian and military customers.



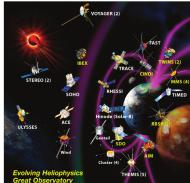
Source: NASA-GSFC

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- Satellites provide remote sensing observations of the Sun and in situ measurements of the solar wind (e.g., ACE, SOHO, STEREO).



Source: NASA-GSFC

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- Satellites provide remote sensing observations of the Sun and in situ measurements of the solar wind (e.g., ACE, SOHO, STEREO).
- The Earth-orbiting spacecraft measure space weather effects in Earth's magnetosphere and ionosphere (e.g., GOES, THEMIS, MMS).



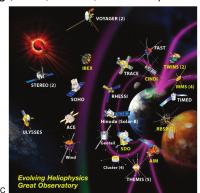
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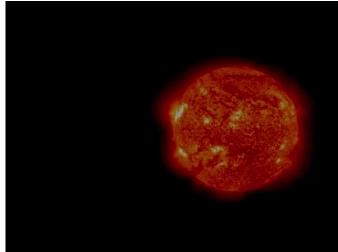
 The Earth-orbiting spacecraft measure space weather effects in Earth's magnetosphere and ionosphere (e.g., GOES, THEMIS, MMS).

 Ground-based observatories provide data for characterizing space weather conditions and effects (INTERMAGNET, SuperMag).



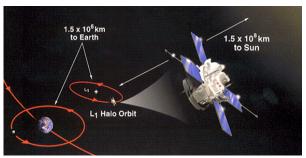
IV. Services Infrastructure: Sun observations

 SOHO measurements allow to predict arrival of MeV protons from solar events that can harm satellites and humans.

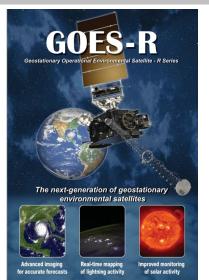


IV. Services Infrastructure: Sun observations

- ACE provides data from its position at Lagrangian point between the Sun and Earth (a point of Earth-Sun gravitational equilibrium about 1.5 million km from Earth and 148.5 million km from the Sun).
- It is a primary data source for measurements of solar particles and magnetic fields.
- ullet ACE provides a \sim 45 minute advance warning before CME strikes Earth.

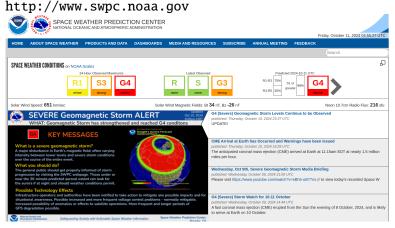


IV. Services Infrastructure: observations from geosynchronous orbit at 6.6 R_E



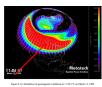
NOAA space weather prediction center products – 50% from GOES (magnetic field, particle fluxes, solar X-ray imager), 38% from ground-based magnetometers, 7% from ground-based solar telescopes, 1% from ACE.

- NOAA uses different scales (G1–G5, S1–S5, R1–R5) to characterize the magnitude and impact of space weather events
- These scales are described in detail on NOAA website:



G-Scale \rightarrow Geomagnetic Storms, Scale G5

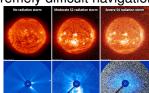
 Power systems: widespread voltage control problems, some grid systems may experience complete collaps or blackouts, transformes may experience damage



- Spacecraft operations: extensive surface charging, problems with orientation
- Other systems: strong pipeline currents, HF radio propagation may be impossible for 1–2 days, degradation of satellite navigation, low-frequency radio navigation can be out for hours
- aurora can be seen as low as Florida and southern Texas (\simeq 40 $^{\circ}$ geo. lat)
- value Kp=9
- 4 days per solar cycle
- Examples: magnetic storms in 1989, 2003 (magnetic storm on Sept 8, 2017, Kp=8, Scale=G4)

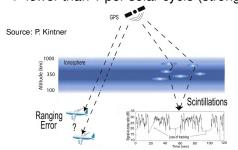
S-Scale → Solar Radiation Storms, Scale S5

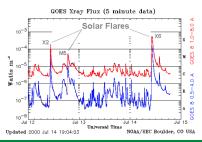
- Biological: unavoidable high radiation hasard to astronauts; passenges and crew in high-flying aircraft at high latitudes may expose radiation risks
- Satellite operations: satellites may be rendered useless, memory impacts can cause loss of control, serios noise in image data, star-trackers may be unable to locate sources, permanent damage to solar panels possible
- Other systems: complete blackout of HF communications possible through polar regions + position errors – extremely difficult navigation
- Flux level of >10 MeV ions
 - Scale S5 ≥10⁵ protons/cm²/sec/ster fewer than 1 per solar cycle
 - * Scale S2 still produces e.g. biological risks ≥10² protons/cm²/sec/ster
 25 per solar cycle



R-Scale → Radio Blackouts. Scale R5

- HF Radio: Complete HF radio blackout on the entire sunlit of the Earth lasting for a number of hours.
 No HF radio contact with mariners and en route aviators
- Navigation: the same but for low frequency signals
- GOES X-ray peak brightness by class X20 and by flux 2⋅10⁻³Watts/m²
- fewer than 1 per solar cycle (strongest X28 on Nov 4, 2003)





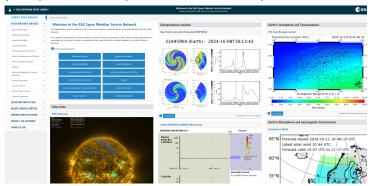
IV. Services Infrastructure: Europe

- Proba-2 mission: EUV imager, STEREO mission
- SWARM mission
- low cost space radiation monitors on as many spacecraft as possible
- many ground-based measurement systems: magnetometers, neutron monitors, GNSS receivers for TEC and ionosondes



IV. Services Infrastructure: Europe

- ESA Space Weather Service Network (https://swe.ssa.esa.int)
- The space weather landscape in Europe is "complicated" and "very fragmented": operational activities from ~25 countries
- competition with other areas of astronomy → limited fundings
- quality of space weather products should be improved

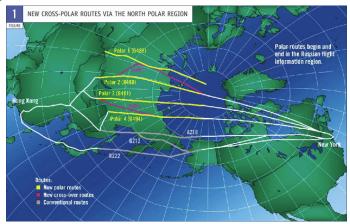


IV. Space Weather Services Infrastructure: Customers

Impact	Customer	Action	Cost
area			
Spacecraft	Lockheed Martin,	Postpone launch,	Loss of
	Boeing, NASA	in orbit reboot sys-	spacecraft
		tems, turn off instru-	\sim \$500M
		ments/spacecraft	
Electric	U.S. Nuclear Reg-	Adjust/reduce system	\$3-6B loss in
Power	ulatory Commis-	load, disconnect com-	GDP (black-
	sion, New York	ponents	out)
	Power Authority		
Airlines	United Airlines,	Divert polar flights,	\sim \$100k per
	Lufthansa, Ko-	change flight plans,	diverted flight
	rean Airlines	change altitude	
Navigation	FAA-WAAS,	Postpone activities, use	\$50k-1M
	Dept. of Trans-	backup systems	daily for sin-
	portation		gle company

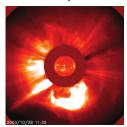
IV. Space Weather Services Infrastructure: Customers

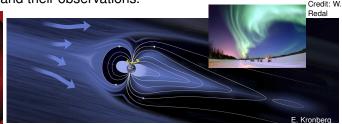
 Using polar routes for air traffic necessitates HF radio communications at high latitudes, which can be disrupted by solar activity.



V. Summary and Outlook

- "It is a challenging task, for both scientific and societal purposes, to develop technologies and mitigation strategies that will help to reliably forecast space weather and its impacts." V. Bothmer
- In the following lectures, we will discuss physical processes that are relevant for space weather: Sun spots, CME, solar flares, high-speed-streams, magnetic storms, substorms, polar lights, ionosphere... and their observations.





VI. Literature

- Severe Space Weather Events Understanding Societatl and Economic Impacts, A Workshop Report, The National Academies Press, Washington, D.C., 2008. DOI: 10.17226/12507
- Walker, C. V. On Magnetic Storms and Earth-Currents. Philos. Trans. R. Soc. Lond. 151, 89–131, 1861.